

15091
Soil
205.5 grams

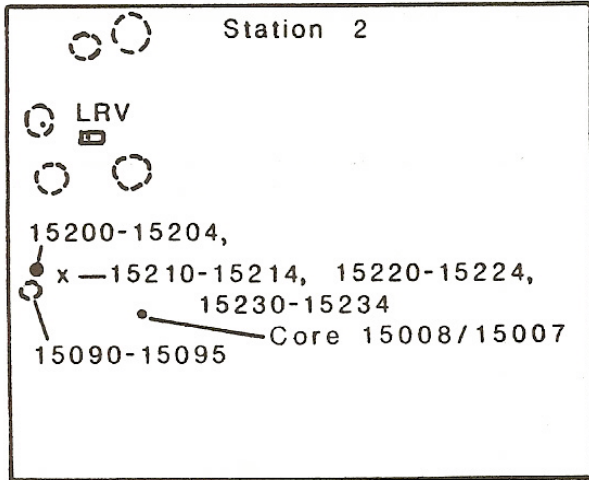


Figure 1: Map of station 2, Apennine Front Apollo 15.



Figure 2: Photo of scoop digging 15090. AS15-86-11549.

Introduction

15091 is a surface soil collected near the boulder at station 2 on the Apennine Front (figures 1 and 2). It should be compared with the top of the double drive tube 15008 collected nearby.

Petrography

The maturity index (I_s/FeO) is 74 (Morris et al. 1978) and it has a high agglutinate content of 53% (Basu et al. 1981). The average grain size is 53 microns (figure 6). It is a mature highland soil with only minor mare component.

Modal content of soil 150291

From Basu et al. 1981.

Agglutinates	55.5%
Basalt	2.8
KREEP basalt	1.8
Breccia	14.7
Anorthosite	1.7
Norite	
Gabbro	0.4
Plagioclase	12.2
Pyroxene	14.2
Olivine	2.8
Ilmenite	0.4
Glass other	10.1

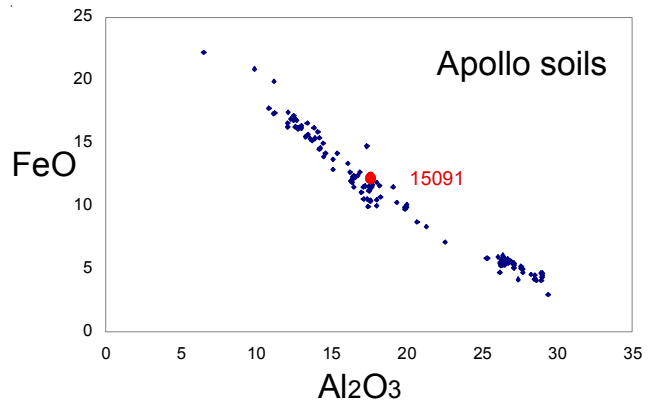


Figure 3: Chemical composition of soil sample 15091 compared with other Apollo soils.

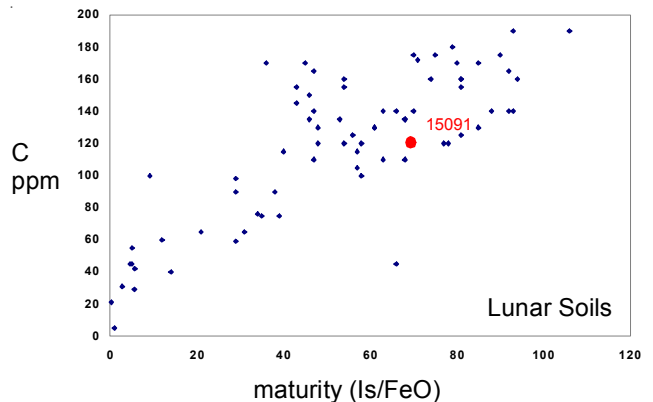


Figure 4: Maturity vrs carbon content inc. 15091.

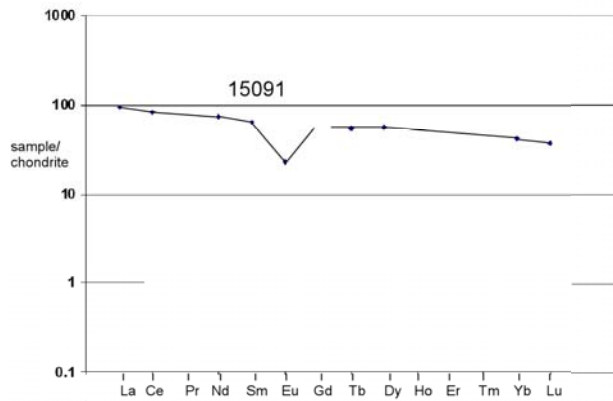


Figure 5: Normalized rare-earth-element diagram for 15091

Chemistry

Fruchter et al. (1973), Wanke et al. (1973) and others analyzed 15091 (figures 3 and 5). Moore et al. (1973) reported 120 ppm C, consistent with maturity (see figure 4). Reed et al. (1972) also determined the halogens, Li, Hg, Te, Ru and Os.

Radiogenic age dating

Murthy et al. (1972) analyzed the Sr isotopic and Rb/Sr ratios.

Cosmogenic isotopes and exposure ages

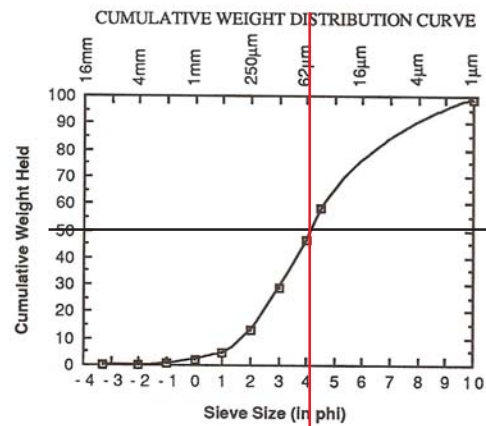
Bhandari et al. (1973) included 15091 in their study of fossil cosmic ray tracks.

Other Studies

Heymann et al. (1972) and Jordan et al. (1974) determined the concentration and isotopic ratio of the rare gasses in 15091.

Processing

15090 was returned in a sealed ALSRC (#1) (Butler 1972). Soil breccias 15095 was returned in same bag.



Average grain size = 53 microns

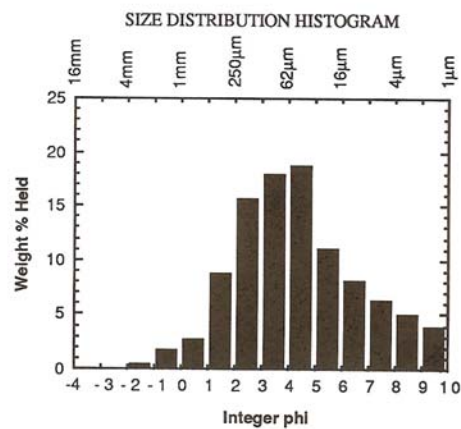
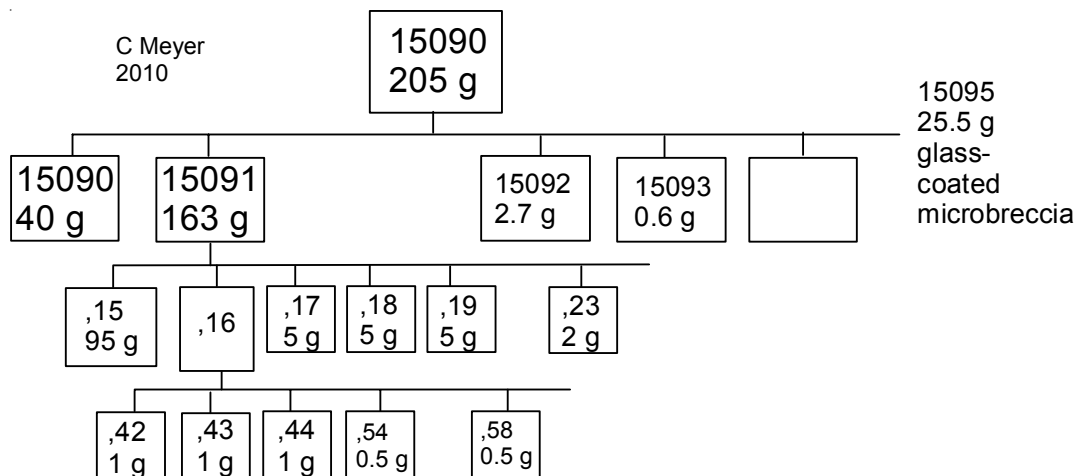


Figure 6: Grain size distribution for 15090 (Graf 1993).

Table 1. Chemical composition of 15091.

<i>reference weight</i>	Fruchter73	Baedecker73	Cuttitta73	Wanke 73	Rancitelli72	Murthy72
SiO ₂ %			46.47	(c) 46.9	(a)	
TiO ₂	1.08	(a)	1.31	(c) 1	(a)	
Al ₂ O ₃	16.6	(a)	17.47	(c) 17.6	(a)	
FeO	11.7	(a)	11.57	(c) 11.75	(a)	
MnO			0.17	(c) 0.15	(a)	
MgO			10.5	(c) 10.4	(a)	
CaO			11.77	(c) 11	(a)	
Na ₂ O	0.41	(a)	0.41	(c) 0.43	(a)	
K ₂ O	0.126	(a)	0.18	(c) 0.17	(a) 0.173	(d) 0.196 (e)
P ₂ O ₅			0.16	(c)		
S %						
<i>sum</i>						
Sc ppm	23	(a)	21	(c) 21	(a)	
V			80	(c)		
Cr	2100	(a)		2120	(a)	
Co	38	(a)	39	(c) 40	(a)	
Ni		251	(b) 365	(c) 170	(a)	
Cu			(b) 7.9	(c)		
Zn		15.2	(b) 16	(c)		
Ga		4.1	(b) 3	(c)		
Ge ppb		410	(b)			
As						
Se						
Rb			5.4	(c)		5.05 (e)
Sr			155	(c)		139.7 (e)
Y			70	(c)		
Zr	250	(a)	240	(c)		
Nb			18	(c)		
Mo						
Ru						
Rh						
Pd ppb						
Ag ppb						
Cd ppb		42	(b)			
In ppb		10.7	(b)			
Sn ppb						
Sb ppb						
Te ppb						
Cs ppm						
Ba	240	(a)	295	(c)		208 (e)
La	21.5	(a)	32	(c) 22	(a)	
Ce	56	(a)		50	(a)	
Pr						
Nd	30	(a)		33	(a)	
Sm	9.8	(a)		9.4	(a)	
Eu	1.34	(a)		1.29	(a)	
Gd						
Tb	1.8	(a)		2	(a)	
Dy				13.6	(a)	
Ho						
Er						
Tm						
Yb	6.7	(a)	7.7	(c) 6.8	(a)	
Lu	1.05	(a)		0.9	(a)	
Hf	7.1	(a)		6.9	(a)	
Ta	1.3	(a)		0.95	(a)	
W ppb						
Re ppb						
Os ppb						
Ir ppb		8.2	(b)			
Pt ppb						
Au ppb		2.7	(b)			
Th ppm	5.8	(a)			3.97	(d)
U ppm					0.93	(d)

technique: (a) INAA, (b) RNAA, (c) "microchemical", (d) radiation counting, (e) IDMS



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